

# Monitoring of EPIC-pn Timing

Jacobo Ebrero

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*Epic Calibration Meeting, 27 April 2017*

# Outline

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- ✧ Report on routine calibration observations of the Crab
  - ✧ Relative timing
  - ✧ Absolute timing
- ✧ Report on the pulse profile anomaly in Timing mode

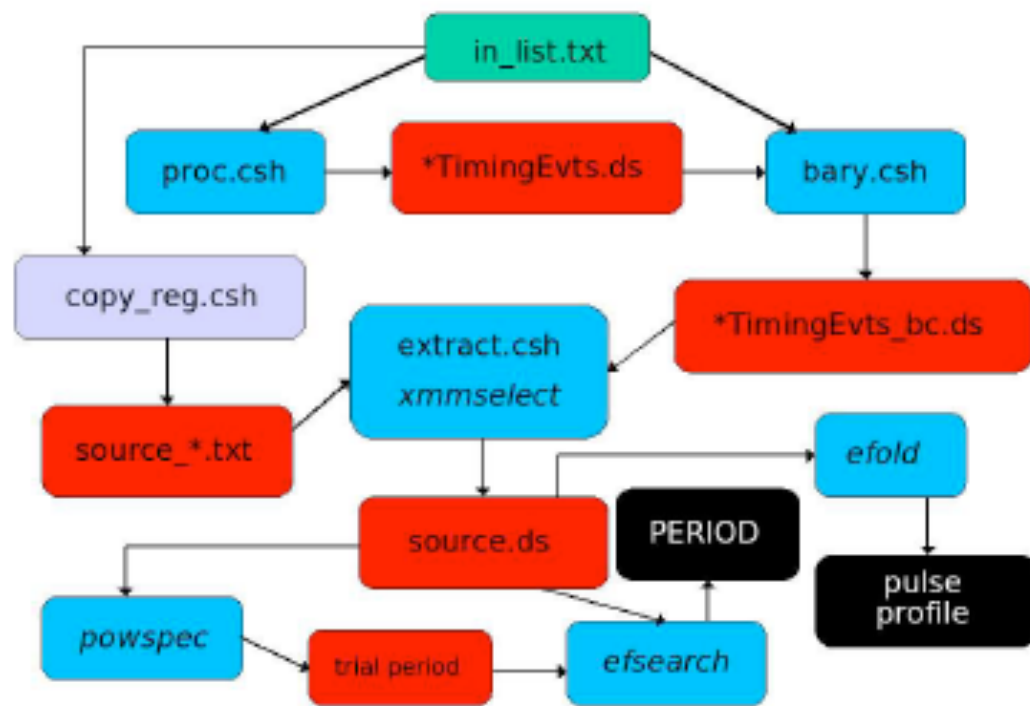


# Relative and absolute timing monitoring

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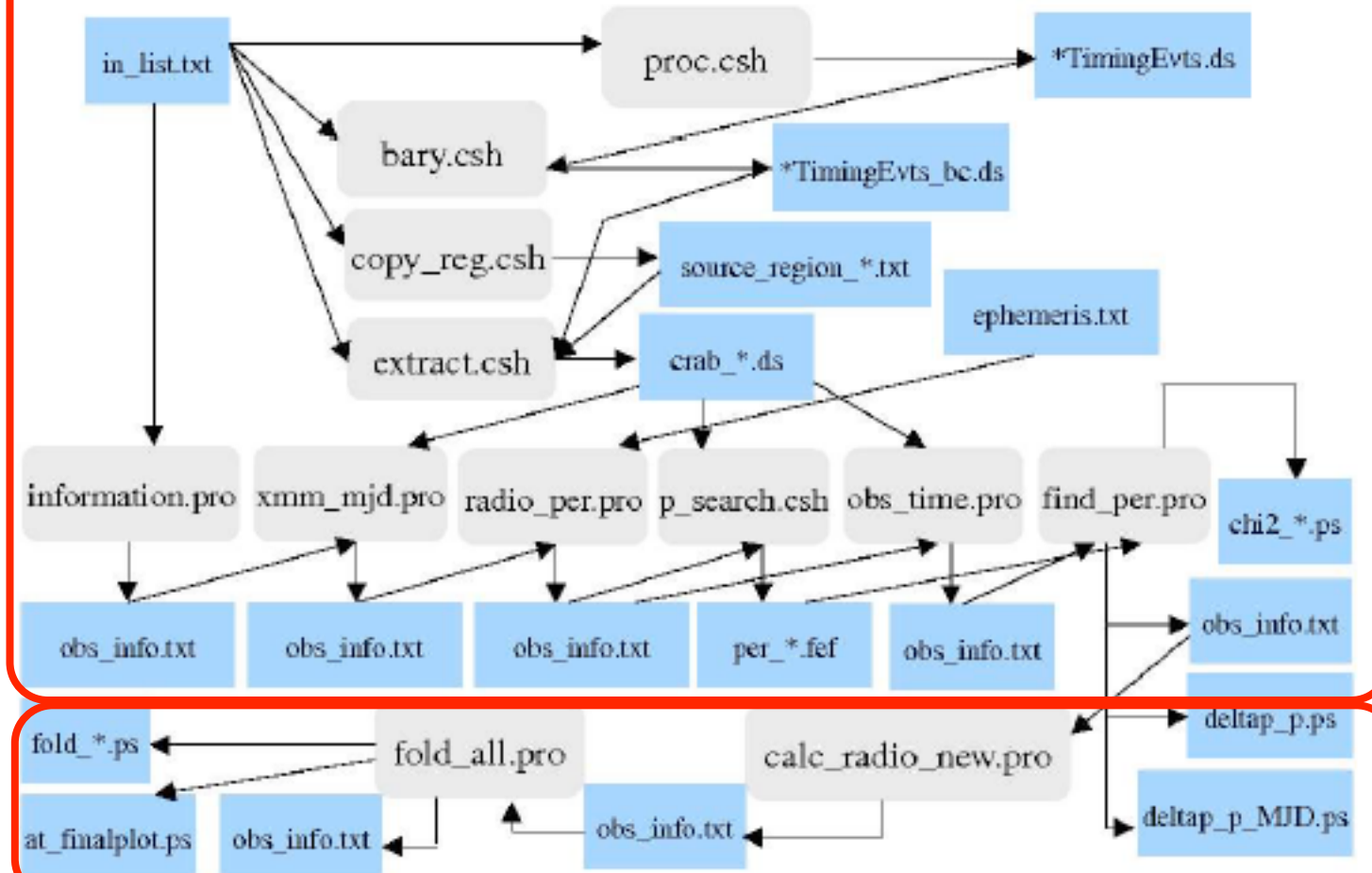
- ❖ *Absolute timing*: locating events in time with reference to standard time defined by atomic clocks or other satellites.
- ❖ *Relative timing*: the capacity to measure time intervals and periodicity reliably.
- ❖ Crab observed twice per semester (spring, autumn).
- ❖  $T_{\text{exp}}$  at least 10 ks, half in Timing and half in Burst mode.
- ❖ Scheduled at different phases of a single orbit to cover different time delays and G/S data links.

# An automated process



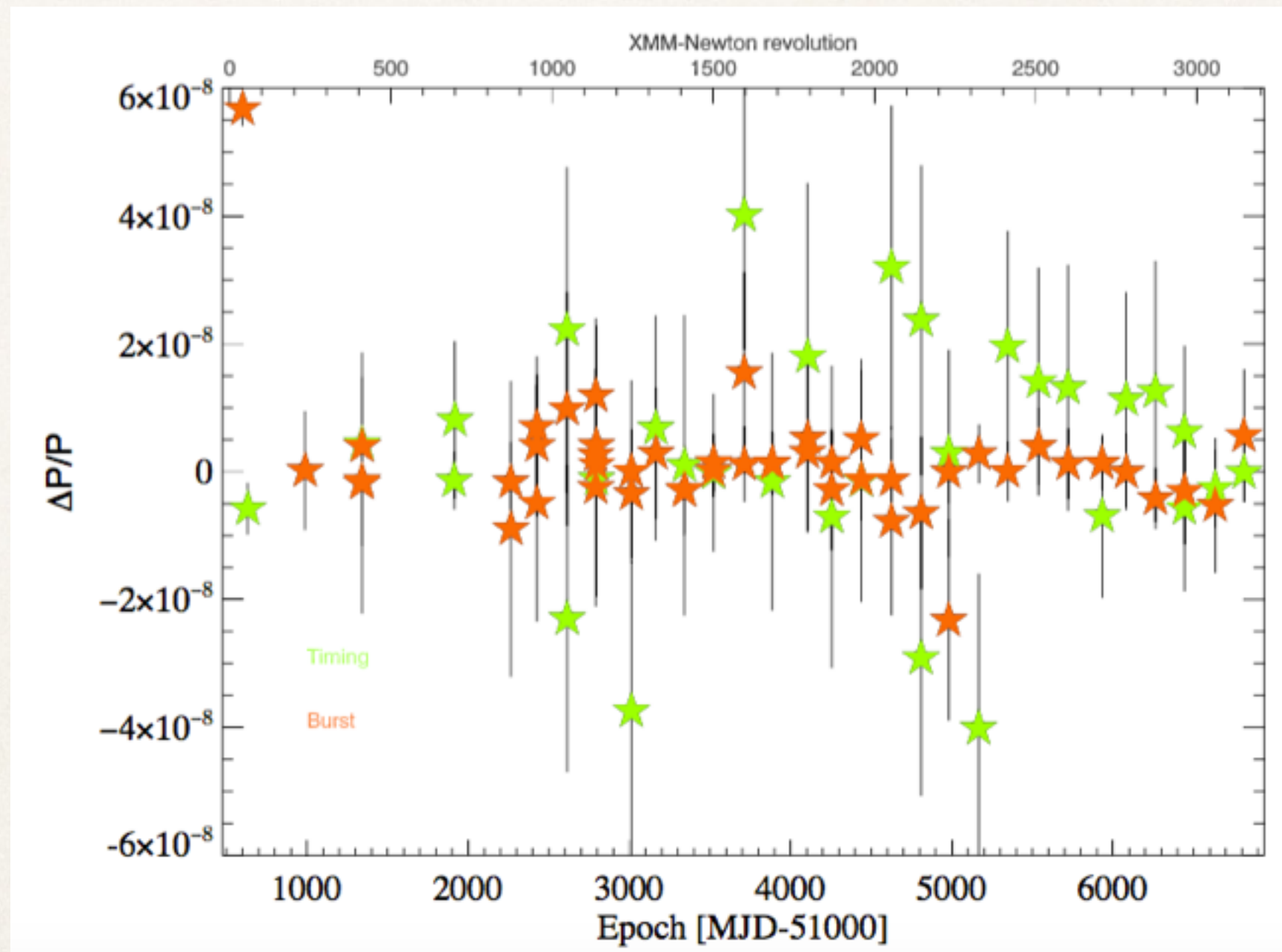
Absolute timing

Relative timing



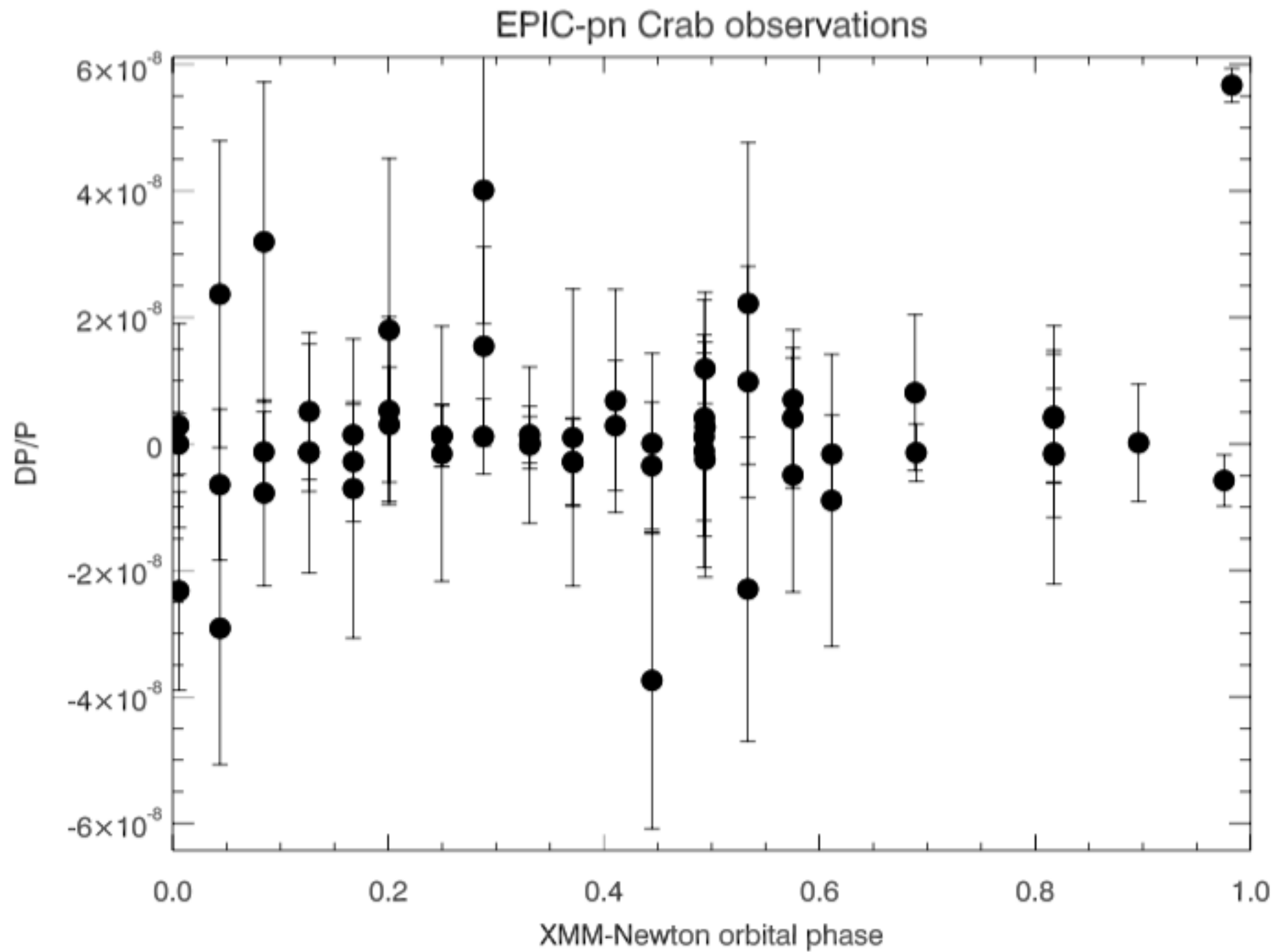


# Relative Timing



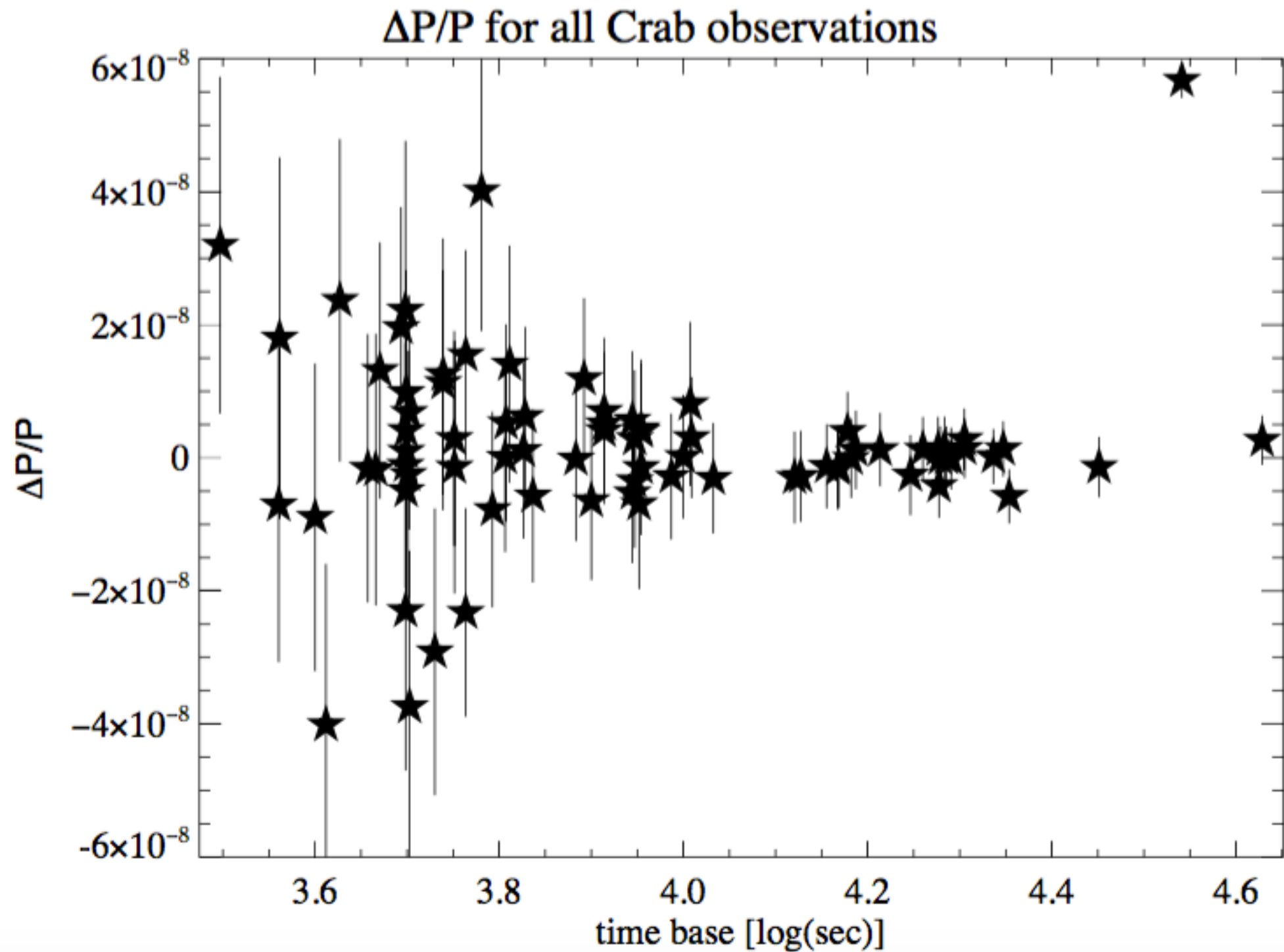
- ❖ Relative deviation of the observed pulse period with respect to the most accurate radio data (Crab ephemeris from Jodrell Bank) is  $< 3 \times 10^{-8}$ .

# Relative Timing

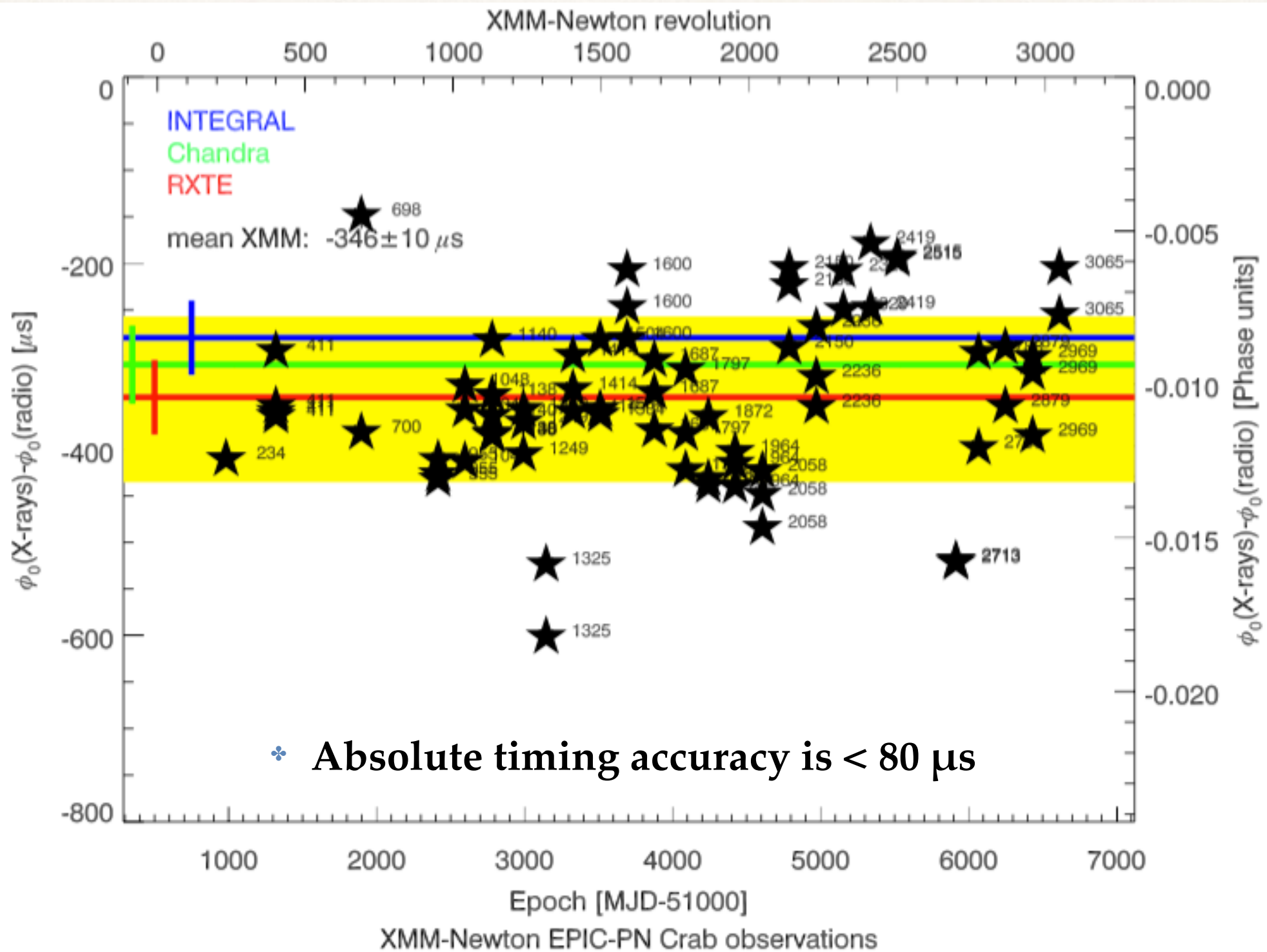




# Relative Timing

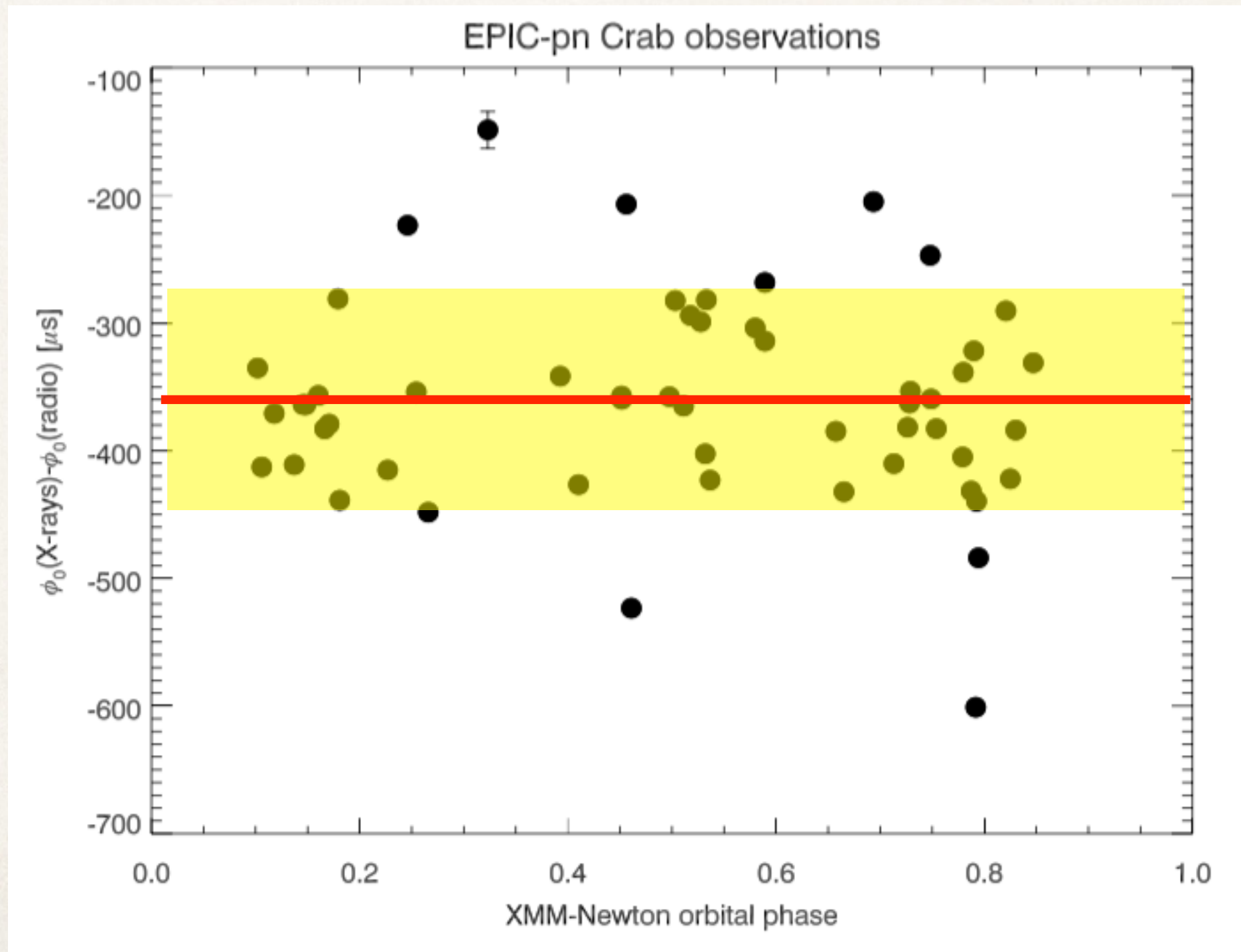


# Absolute Timing





# Absolute Timing



# Pulse profile anomaly in Timing mode

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## *XMM-Newton* Calibration Technical Note

Seasonal pulse profile distortions in the EPIC-pn timing  
observations of the Crab pulsar

J. Ebrero, M. J. S. Smith (ESAC)  
M. J. Freyberg (MPE Garching)

March 24, 2017



# Pulse profile anomaly in Timing mode

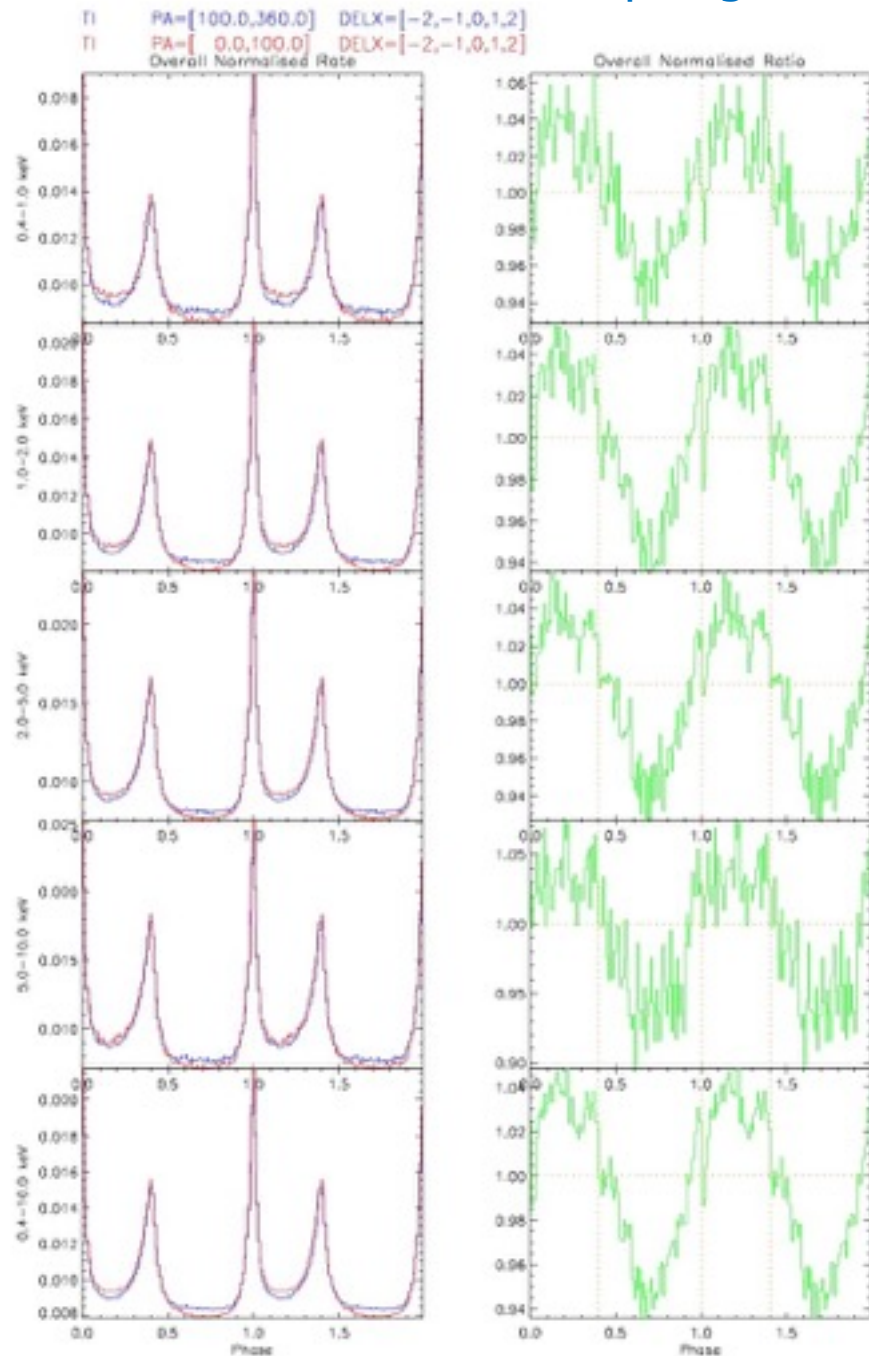
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- ❖ The pulse profile of the Crab in TI mode show systematic differences between Spring and Autumn observations. BU mode observations are unaffected.
- ❖ Spring TI mode profiles are similar to those of BU mode observations. Autumn TI mode profiles are distorted (excess / deficit of counts in the inter-pulse valleys).
- ❖ This effect has been present since beginning of mission.



# Pulse profile anomaly in Timing mode

TI Autumn / TI Spring



0.4-1.0 keV

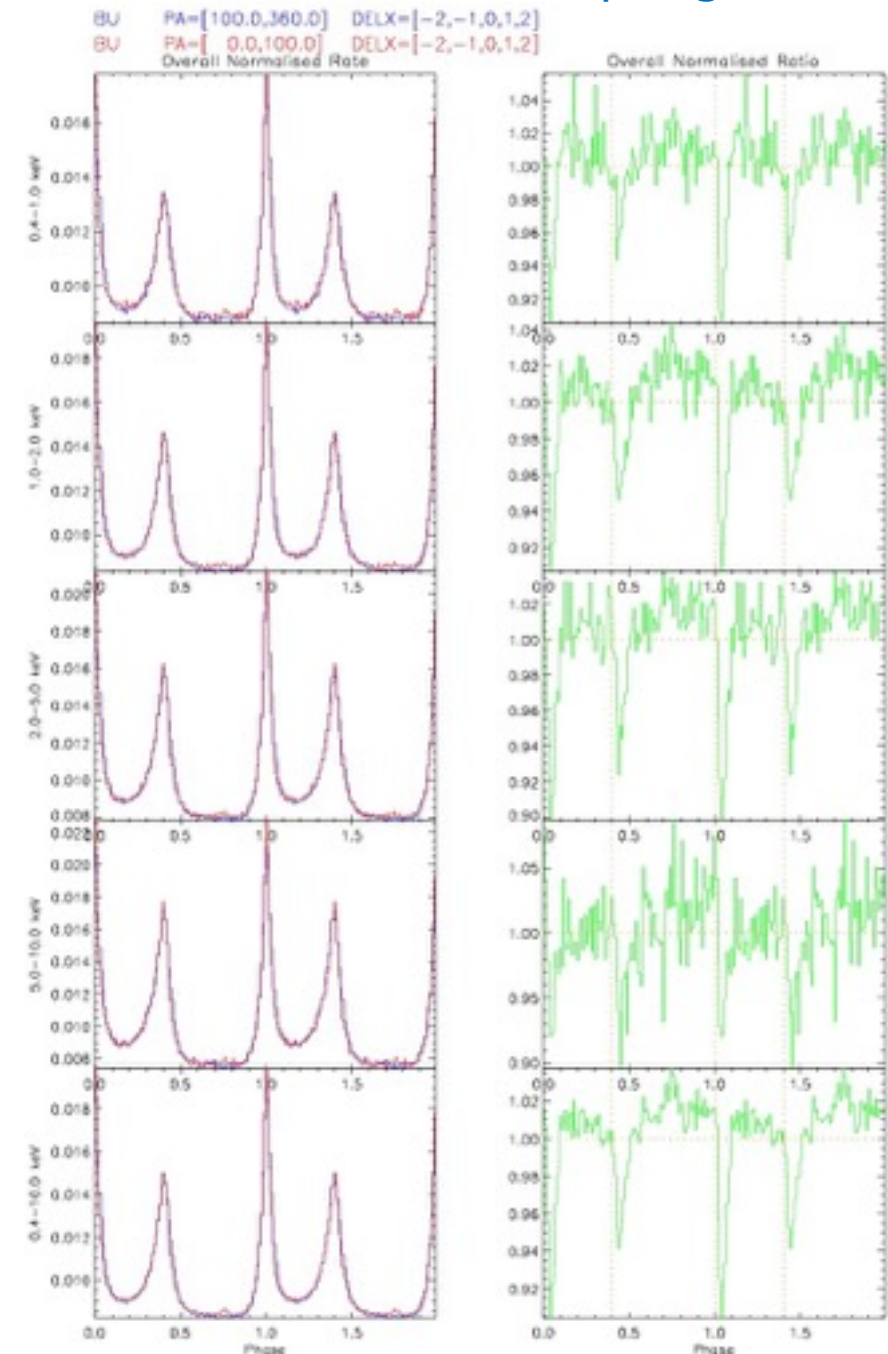
1.0-2.0 keV

2.0-5.0 keV

5.0-10 keV

0.4-10 keV

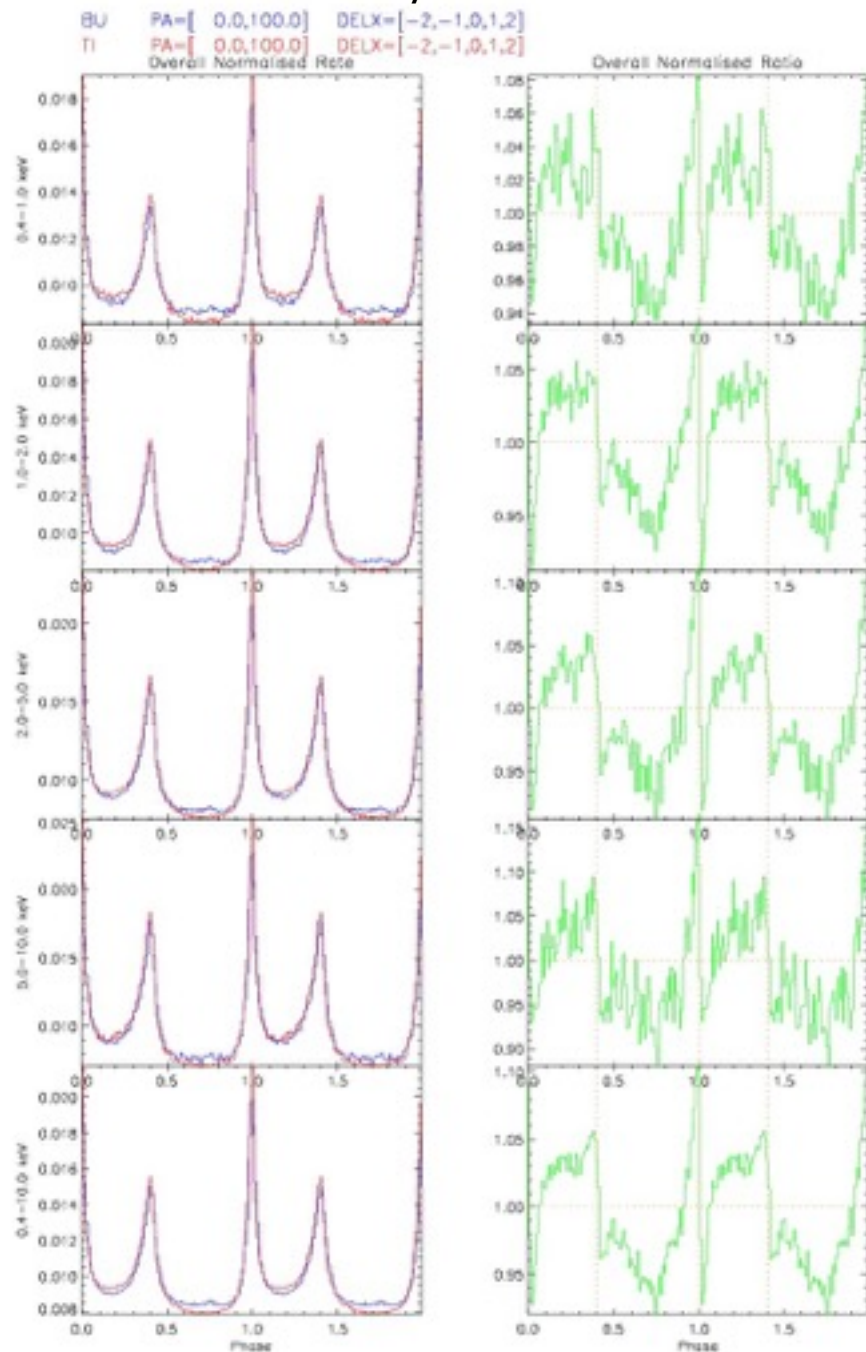
BU Autumn / BU Spring





# Pulse profile anomaly in Timing mode

TI Autumn / BU Autumn



0.4-1.0 keV

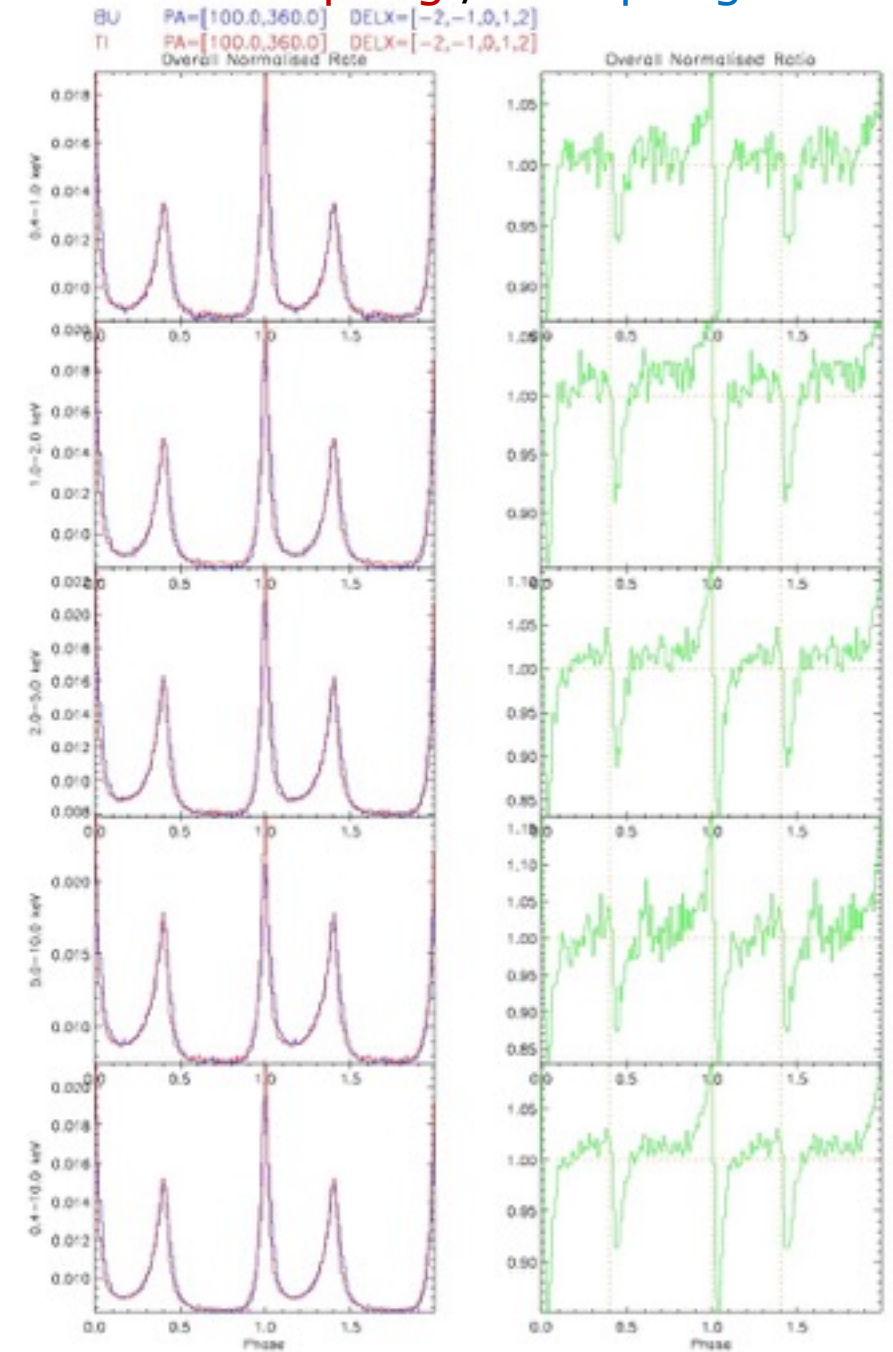
1.0-2.0 keV

2.0-5.0 keV

5.0-10 keV

0.4-10 keV

TI Spring / BU Spring





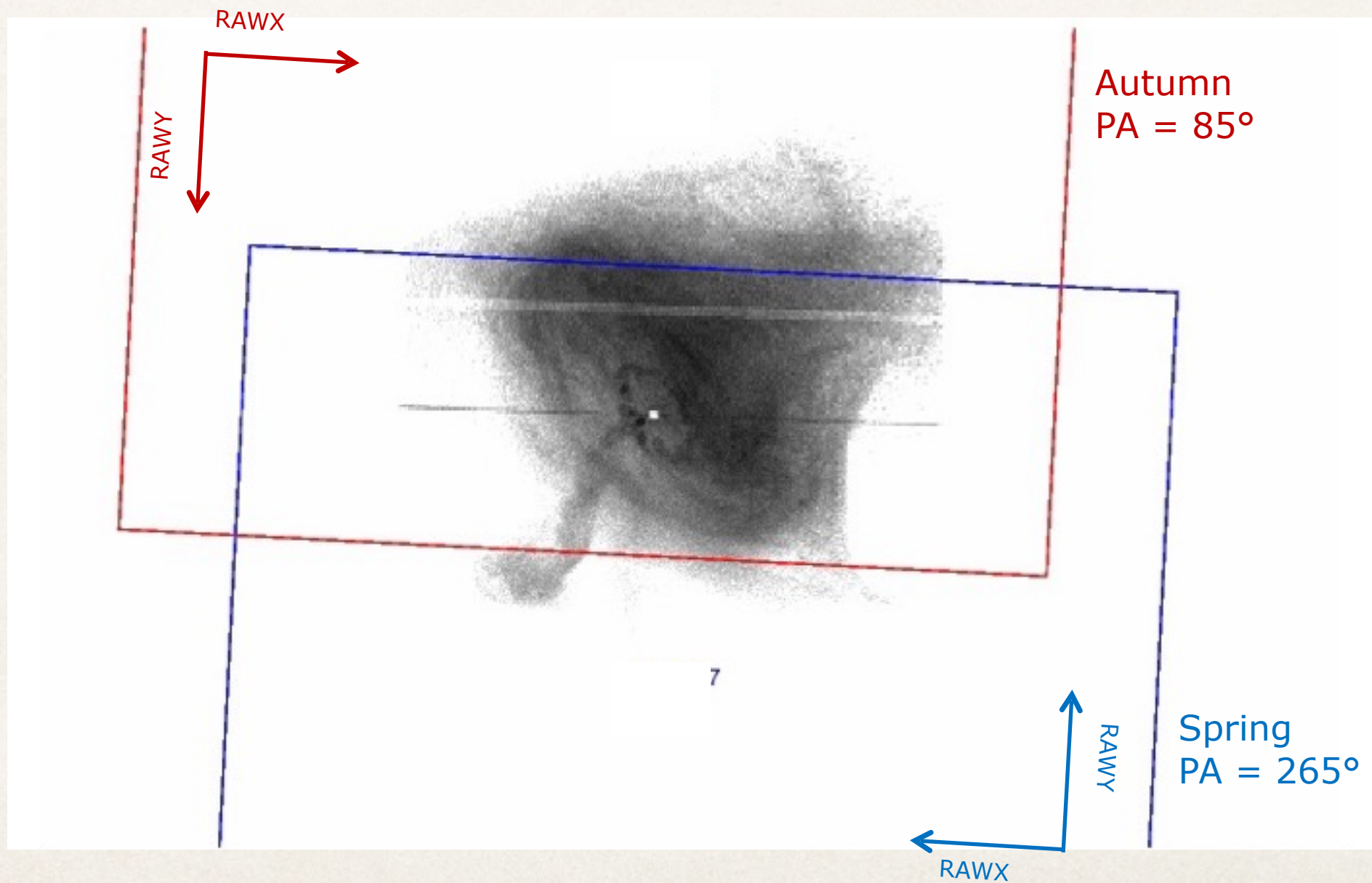
# Pulse profile anomaly in Timing mode

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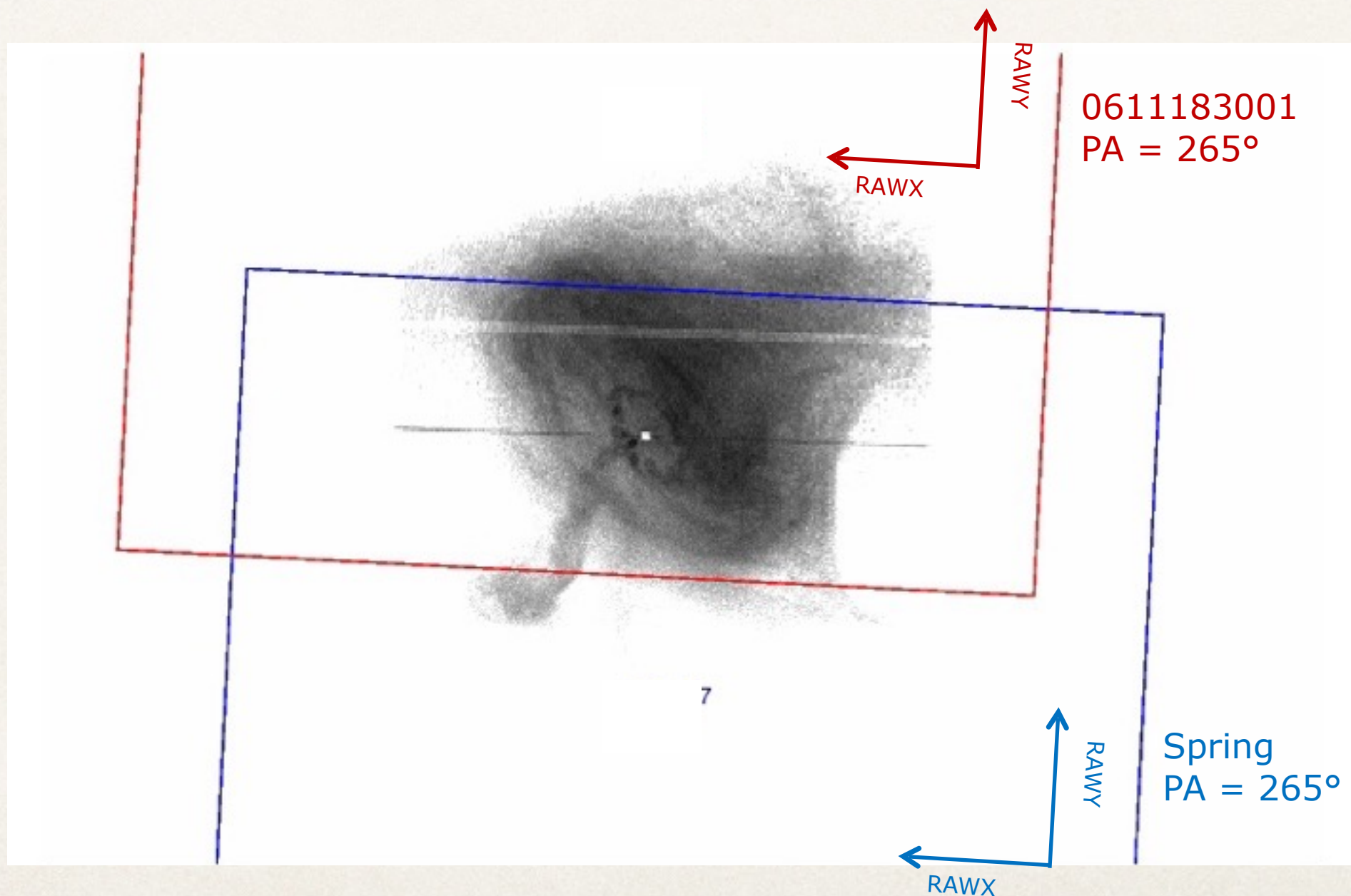
- ❖ Investigation of the anomaly:
  - ❖ Change in position angle



# Pulse profile anomaly in Timing mode



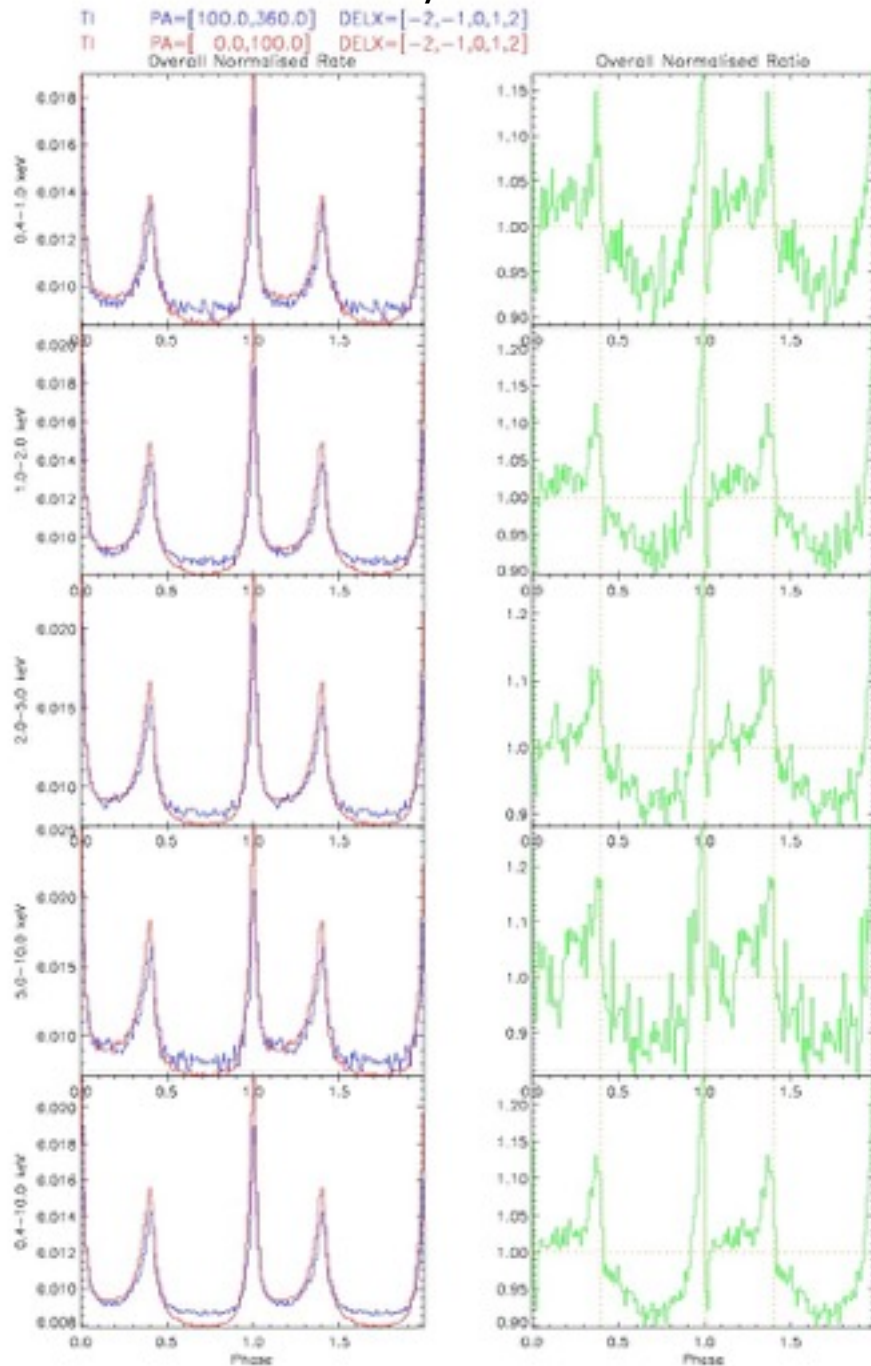
# Pulse profile anomaly in Timing mode





# Pulse profile anomaly in Timing mode

TI Autumn / 0611183001



0.4-1.0 keV

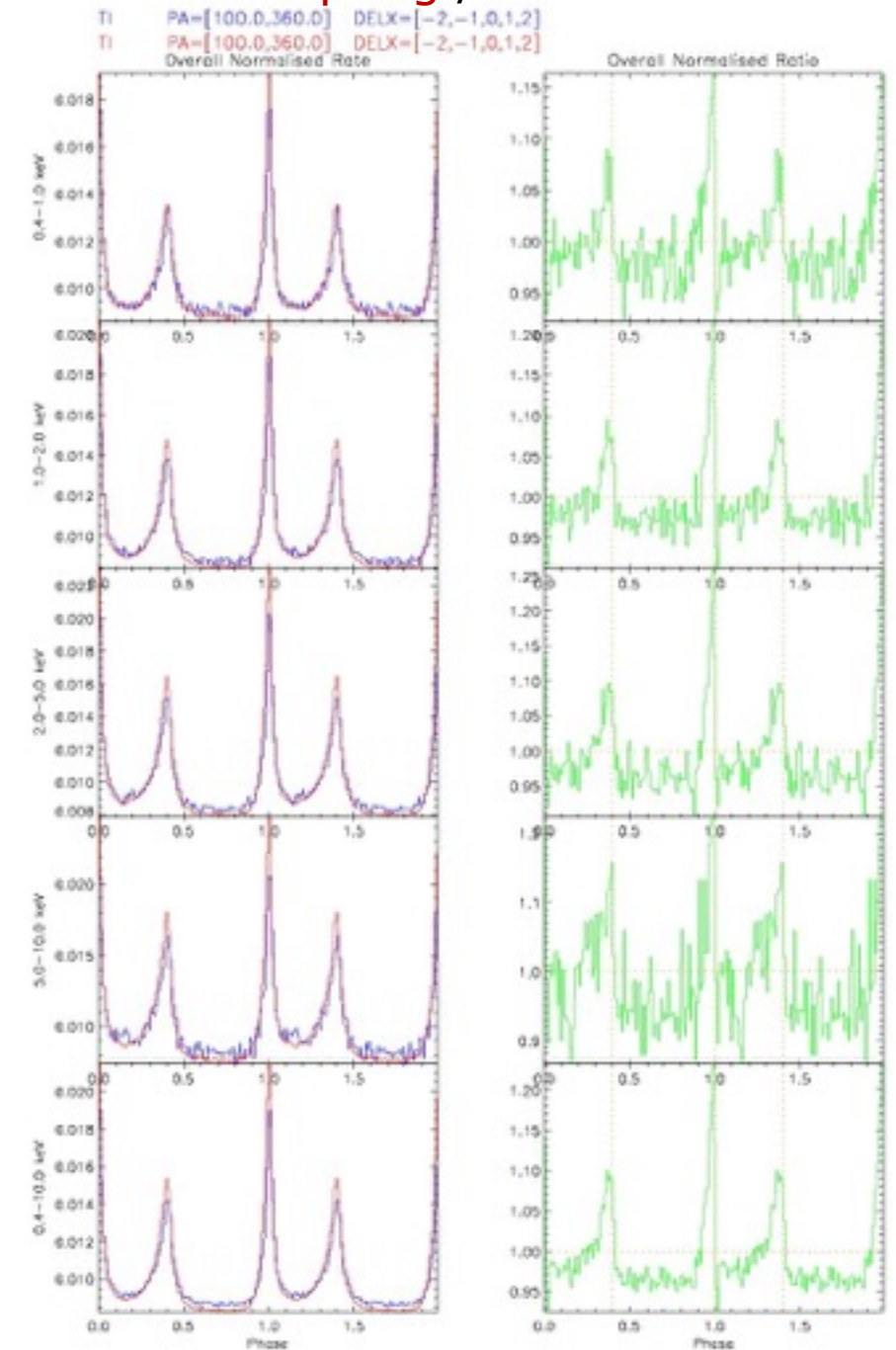
1.0-2.0 keV

2.0-5.0 keV

5.0-10 keV

0.4-10 keV

TI Spring / 0611183001





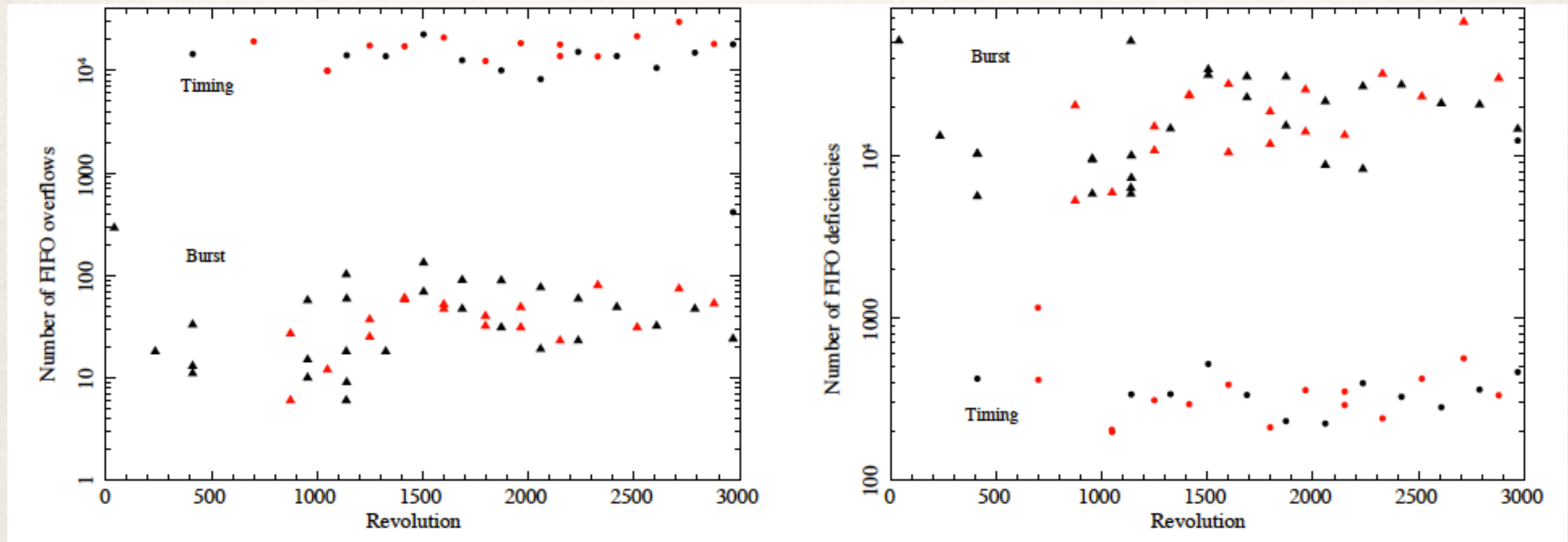
# Pulse profile anomaly in Timing mode

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- ❖ Investigation of the anomaly:
  - ❖ Change in position angle
  - ❖ Housekeeping parameters

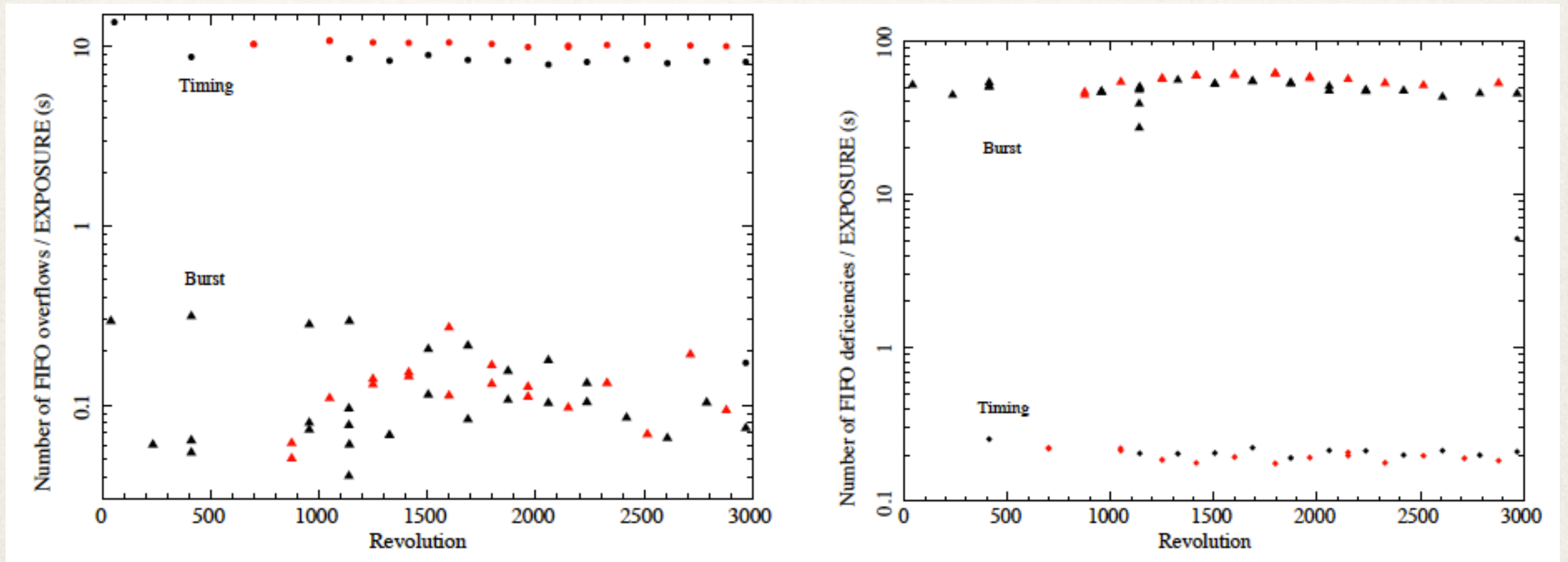


# Pulse profile anomaly in Timing mode



- ❖ Number of FIFO overflows much higher in TI than in BU mode.
- ❖ Apparently, no seasonal effect observed...

# Pulse profile anomaly in Timing mode



- ❖ Normalized number of FIFO overflows is systematically higher in Autumn in TI observations. No trend is observed for BU mode observations.



# Pulse profile anomaly in Timing mode

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- ❖ Preliminary conclusions:

- ❖ FIFO overflows cause a loss of counts at different phases of the Crab pulse profile.
- ❖ Seasonal dependence is due to the different number of counts gathered on-board because of the different coverage of the nebula.
- ❖ For very bright sources do not use TI, but rather BU.